Clifford Space as a Generalization of Spacetime: Prospects for QFT of Point Particles and Strings

Matej Pavšič Jožef Stefan Institute, Jamova 39, 1000 Ljubljana, Slovenia e-mail: matej.pavsic@ijs.si

Abstract

The idea that spacetime has to be replaced by Clifford space (C-space) is explored. Quantum field theory (QFT) and string theory are generalized to C-space. It is shown how one can solve the cosmological constant problem and formulate string theory without central terms in the Virasoro algebra by exploiting the peculiar pseudo-Euclidean signature of C-space and the Jackiw definition of the vacuum state. As an introduction into the subject, a toy model of the harmonic oscillator in pseudo-Euclidean space is studied.

1 Introduction

Quantum field theory is a very successful theory, but also enigmatic. There occur infinities which require renormalization. This suggests that the theory is not yet complete¹. An approach whose roots go back to Feynman² [2] employs an invariant evolution parameter τ , introduced by Fock and Stueckelberg [4], and subsequently pursued by many authors [5, 6]. Another direction of research starts from the idea

¹P.A.M. Dirac expressed such a view, e.g., in a talk presented at the Erice 1982 conference [1]. ²See a nice paper by Schweber [3].

that spacetime has to be replaced by a more general space, namely the Clifford space (shortly, C-space) [6]–[13]. It has been found [9, 6] that the Stueckelberg theory with the Lorentz invariant evolution parameter τ naturally occurs as being embedded in the theory based on Clifford space. In this paper we would like to point out another important aspect of Clifford space. We will show that quantum field theory generalized to Clifford space provides a natural way of resolving the notorious cosmological constant problem. We exploit the property of Clifford space with signature (+++...---...), where the number of plus and minus signs is the same, provided that the underlying spacetime has Minkowski signature. We find that by using the Jackiw definition of vacuum [14], the concept of C-space enables a formulation of QFT in which zero point energies belonging to positive and negative signature degrees of freedom cancel out, while preserving the Casimir effect.

Introduction of C-space has consequences for string theory which can be formulated without central terms in Virasoro algebra even when the dimension of the underlying spacetime is four. We do not need a higher dimensional target spacetime for a consistent formulation of (quantized) string theory. Instead of a higher dimensional space we have the 16-dimensional Clifford space which also provides a natural framework for description of superstings and supersymmetry, since spinors are just the elements of left or right minimal ideals of Clifford algebra [16]–[18], [6, 25].

After a brief review of the concept of Clifford space we first discuss a toy model of the harmonic oscillator in pseudo-Euclidean space. We employ the obvious fact that if a Lagrangian is multiplied by -1, whilst the definitions of momentum and energy are kept the same, namely, $p_{\mu} = \partial L/\dot{x}^{\mu}$, $E = p_{\mu}\dot{x}^{\mu} - L$, then energy becomes negative. In such case the criterion for stability is reversed: the system is stable when its energy has maximum. This is precisely what happens in a pseudo-Euclidean space: The Lagrangian is a quadratic form which has the terms with positive and negative signs. Therefore the expression for energy is also composed of the terms with positive and negative signs.

We then describe a system of n scalar fields forming a space $M_{r,-s}$, n = r + s, and show that when r = s, the zero point energy vanishes. The vacuum contribution to the stress-energy tensor is zero, and there is no cosmological constant problem [26]

in this model³. Then we discuss a model in which the space of n-fields is generalized to the corresponding Clifford space.

Finally, we discuss the string theory and generalize it to C-space.

2 Clifford Space as the Arena for Physics

Clifford algebra is a very useful tool for description of geometry. Since Hestenes's seminal books [17], there is a growing interest in using Clifford algebra in physics (see e.g. refs. [18]–[24]), and generalizing physics from spacetime to Clifford space [6]–[13]. An n-dimensional flat space (e.g., spacetime) M_n can be described by means of a complete set of basis vectors γ_{μ} , $\mu = 0, 1, 2, ..., n-1$, which satisfy the Clifford algebra relations

$$\gamma_{\mu} \cdot \gamma_{\nu} \equiv \frac{1}{2} (\gamma_{\mu} \gamma_{\nu} + \gamma_{\nu} \gamma_{\mu}) = g_{\mu\nu} \tag{1}$$

This is just the symmetric part of the Clifford (or geometric) product $\gamma_{\mu}\gamma_{\nu}$. It is equal to the metric $g_{\mu\nu}$. The antisymmetric part defines a bivector which represents an oriented unit area:

$$\gamma_{\mu} \wedge \gamma_{\nu} = \frac{1}{2} (\gamma_{\mu} \gamma_{\nu} - \gamma_{\nu} \gamma_{\mu}) \equiv \frac{1}{2} [\gamma_{\mu}, \gamma_{\nu}]$$
 (2)

This can be continued to the antisymmetrized product of 3, 4, ..., n basis vectors

$$\gamma_{\mu_1} \wedge \gamma_{\mu_2} \wedge \gamma_{\mu_3} = \frac{1}{3!} [\gamma_{\mu_1}, \gamma_{\mu_2}, \gamma_{\mu_3}] \tag{3}$$

$$\vdots$$
 (4)

$$\gamma_{\mu_1} \wedge \gamma_{\mu_2} \wedge \dots \wedge \gamma_{\mu_n} = \frac{1}{r!} [\gamma_{\mu_1}, \gamma_{\mu_2}, \dots, \gamma_{\mu_n}]$$

$$(5)$$

An object $\gamma_{\mu_1} \wedge \gamma_{\mu_2} \wedge ... \wedge \gamma_{\mu_r}$ of degree $r, 1 \leq r \leq n$, is called a basis *multivector* or r-vector. it represents an oriented r-dimensional unit area.

A point P in M_n can be associated with a *vector* x joining the coordinate origin O and P:

$$x = x^{\mu} \gamma_{\mu} \tag{6}$$

³An eventual small non-vanishing cosmological constant, as confirmed by recent observations [27], can be a residual effect of something else, or due to our incomplete understanding of the dynamical laws at cosmological scales.

A generic object is a Clifford number, called also a *polyvector* or *Clifford aggregate*, which is a superposition of multivectors⁴:

$$X = \sigma \underline{1} + x^{\mu} \gamma_{\mu} + \frac{1}{2} x^{\mu_{1} \mu_{2}} \gamma_{\mu_{1} \mu_{1}} + \dots + \frac{1}{n!} x^{\mu_{1} \dots \mu_{n}} \gamma_{\mu_{1} \dots \mu_{n}} \equiv x^{M} \gamma_{M}$$
 (7)

Here $\gamma_{\mu_1...\mu_r} \equiv \gamma_{\mu_1} \wedge \gamma_{\mu_2} \wedge ... \wedge \gamma_{\mu_r}$ and

$$x^{M} = (\sigma, x^{\mu}, x^{\mu_{1}\mu_{2}}, ..., x^{\mu_{1}...\mu_{r}})$$

$$\gamma_{M} = (\underline{1}, \gamma_{\mu}, \gamma_{\mu_{1}\mu_{2}}, ..., \gamma_{\mu_{1}...\mu_{r}}), \quad \mu_{1} < \mu_{2} < ... < \mu_{r}$$
(8)

are respectively coordinates and basis elements of Clifford algebra.

The coordinates $X^{\mu_1...\mu_r}$ determine an oriented r-area. They say nothing about the precise form of the (r-1)-loop enclosing the r-area. The coordinates σ , x^{μ} , $x^{\mu_1\mu_2}$,... provide a means for a description of extended objects. If an object is extended, then not only its center of mass coordinates x^{μ} , but also the higher grade coordinates $x^{\mu_1\mu_2}$, $x^{\mu_1\mu_2,\mu_3}$,..., associated with the object extension, are different from zero, in general. Those higher grade coordinates model the extended object. They are a generalization of the concept of center of mass [10].

Since x^M assumes any real value, the set of all possible X forms a 2^n -dimensional manifold, called *Clifford space*, or shortly C-space.

Let us define the quadratic form by means of the scalar product

$$|dX|^2 \equiv dX^{\ddagger} * dX = dx^M dx^N G_{MN} \equiv dx^M dx_M$$
(9)

where the metric of C-space is given by

$$G_{MN} = \gamma_M^{\ddagger} * \gamma_N \tag{10}$$

The operation ‡ reverses the order of vectors:

$$(\gamma_{\mu_1}\gamma_{\mu_2}...\gamma_{\mu_r})^{\dagger} = \gamma_{\mu_r}...\gamma_{\mu_2}\gamma_{\mu_1} \tag{11}$$

Indices are lowered and raised by G_{MN} and its inverse G^{MN} , respectively. The following relation is satisfied:

$$G^{MJ}G_{JN} = \delta^{M}{}_{N} \tag{12}$$

⁴Although Hestenes and others use the term 'multivector' for a generic Clifford number, we prefer to call it 'polyvector', and reserve the name 'multivector' for objects of definite grade. So our nomenclature is in agreement with the one used in the theory of differential forms, where 'multivectors' mean objects of definite grade, and not a superposition of objects with different grade.

Considering the definition (10) for the C-space metric, one could ask why just that definition, which involves reversion, and not a slightly different definition, e.g., without reversion. That reversion is necessary for consistency we can demonstrate by the following example. Let us take a polyvector which has only the 2-vector component different from zero:

$$x^{N} = (0, 0, x^{\alpha\beta}, 0, 0, ..., 0)$$
(13)

Then the covariant components are

$$x_M = G_{MN} x^N = \frac{1}{2} G_{M[\alpha\beta]} x^{\alpha\beta} \tag{14}$$

Since the metric G_{MN} is block diagonal, so that $G_{M[\alpha\beta]}$ differs from zero only if M is bivector index, we have

$$x_M = x_{\mu\nu} = \frac{1}{2} G_{[\mu\nu][\alpha\beta]} x^{\alpha\beta} \tag{15}$$

From the definition (10) we find

$$G_{[\mu\nu][\alpha\beta]} = (\gamma_{\mu} \wedge \gamma_{\nu})^{\ddagger} * (\gamma_{\alpha} \wedge \gamma_{\beta}) = (\gamma_{\nu} \wedge \gamma_{\mu}) * (\gamma_{\alpha} \wedge \gamma_{\beta}) = g_{\mu\alpha}g_{\nu\beta} - g_{\mu\beta}g_{\nu\alpha}$$
 (16)

Inserting (16) into (15) we obtain

$$x_{\mu\nu} = \frac{1}{2} (g_{\mu\alpha}g_{\nu\beta} - g_{\mu\beta}g_{\nu\alpha})x^{\alpha\beta} = g_{\mu\alpha}g_{\nu\beta}x^{\alpha\beta}$$
 (17)

From the fact that the usual metric $g_{\mu\nu}$ lowers the indices μ , ν , α , β ,..., so that

$$g_{\mu\alpha}g_{\nu\beta}x^{\alpha\beta} = x_{\mu\nu} \tag{18}$$

It follows that eq. (17) is just an identity.

Had we defined the C-space metric without employing the reversion, then instead of eq.(16) and (17) we would have $G_{[\mu\nu][\alpha\beta]} = -(g_{\mu\alpha}g_{\nu\beta} - g_{\mu\beta}g_{\nu\alpha})$ and $x_{\mu\nu} = -g_{\mu\alpha}g_{\nu\beta}x^{\alpha\beta} = -x_{\mu\nu}$, which is a contradiction⁵.

Eq.(9) is the expression for the *line element* in C-space. If C-space is generated from the basis vectors γ_{μ} of spacetime M_n with signature (+ - - - - ...), then

⁵By an analogous derivation we find that the relation $x^M = G^M{}_N x^N = \delta^M{}_N x^N$ holds if $G^M{}_N = (\gamma^M)^{\ddagger} * \gamma_N$. The definition $G^M{}_N = \gamma^M * \gamma_N$ leads to the contradictory equation $x^M = G^M{}_N x^N = -\delta^M{}_N x^N$.

the signature of C-space is (+++...--...), where the number of plus and minus signs is the same, namely, $2^n/2$. This has some important consequences that we are going to investigate in the next sections.

We assume that 2^n -dimensional Clifford space is the arena in which physics takes place. We can take n = 4, so that the spacetime from which we start is just the 4-dimensional Minkowski space M_4 . The corresponding Clifford space has then 16 dimensions. In C-space the usual points, lines, surfaces, volumes and 4-volumes are all described on the same footing and can be transformed into each other by rotations in C-space (called polydimensional rotations):

$$x'^{M} = L^{M}{}_{N}x^{N} \tag{19}$$

subjected to the condition $|dX'|^2 = |dX|^2$.

We can now envisage that physical objects are described in terms of $x^M = (\sigma, x^{\mu}, x^{\mu\nu}, ...)$. The first straightforward possibility is to introduce a single parameter τ and consider a mapping

$$\tau \to x^M = X^M(\tau) \tag{20}$$

where $X^M(\tau)$ are 16 embedding functions that describe a world-line in C-space. From the point of view of C-space, $X^M(\tau)$ describe a wordline of a "point particle": At every value of τ we have a point in C-space. But from the perspective of the underlying 4-dimensional spacetime, $X^M(\tau)$ describe an extended object, sampled by the center of mass coordinates $X^{\mu}(\tau)$ and the coordinates $X^{\mu_1\mu_2}(\tau), ..., X^{\mu_1\mu_2\mu_3\mu_4}(\tau)$. They are a generalization of the center of mass coordinates in the sense that they provide information about the object 2-vector, 3-vector, and 4-vector extension and orientation⁶

Let the dynamics of such an object be determined by the action

$$I[X] = M \int d\tau \, (\dot{X}^{\ddagger} * \dot{X})^{\frac{1}{2}} = M \int d\tau (\dot{X}^{M} \dot{X}_{M})^{\frac{1}{2}}$$
 (21)

The dynamical variables are given by the polyvector

$$X = X^{M} \gamma_{M} = \sigma \underline{1} + X^{\mu} \gamma_{\mu} + X^{\mu_{1} \mu_{2}} \gamma_{\mu_{1} \mu_{2}} + ... X^{\mu_{1} ... \mu_{n}} \gamma_{\mu_{1} ... \mu_{n}}$$
(22)

⁶A systematic and detailed treatment is in ref. [10].

whilst

$$\dot{X} = \dot{X}^{M} \gamma_{M} = \dot{\sigma} \underline{1} + \dot{X}^{\mu} \gamma_{\mu} + \dot{X}^{\mu_{1}\mu_{2}} \gamma_{\mu_{1}\mu_{2}} + \dots \dot{X}^{\mu_{1}\dots\mu_{n}} \gamma_{\mu_{1}\dots\mu_{n}}$$
(23)

is the velocity polyvector, where $\dot{X}^M \equiv dX^M/d\tau$.

In the action (21) we have a straightforward generalization of the relativistic point particle in M_4 :

$$I[X^{\mu}] = m \int d\tau (\dot{X}^{\mu} \dot{X}_{\mu})^{\frac{1}{2}} , \quad \mu = 0, 1, 2, 3$$
 (24)

If a particle is extended, then the latter action (24) provides only a very incomplete description. A more complete description is given by the action (21), in which the C-space embedding functions $X^M(\tau)$ sample the objects extension [10].

3 A toy model: Harmonic oscillator in pseudo-Eucidean space

3.1 A simple model in $M_{1,-1}$

Suppose that instead of usual Harmonic oscillator we have a system given by the Lagrangian [15]

$$L = \frac{1}{2}(\dot{x}^2 - \dot{y}^2) - \frac{1}{2}\omega^2(x^2 - y^2)$$
 (25)

If we derive the equations of motion we find that they are indistinguishable form those of the usual harmonic oscillator:

$$\ddot{x} + \omega^2 x = 0 , \qquad \ddot{y} + \omega^2 y = 0$$
 (26)

The change of sign in front of the y-term have no influence on the equations of motion.

However, a difference occurs when we calculate the *canonical momenta*

$$p_x = \frac{\partial L}{\partial \dot{x}} = \dot{x} , \qquad p_y = \frac{\partial L}{\partial \dot{y}} = -\dot{y}$$
 (27)

and the Hamiltonian

$$H = p_x \dot{x} + p_y \dot{y} - L = \frac{1}{2} (p_x^2 - p_y^2) + \frac{\omega^2}{2} (x^2 - y^2)$$
 (28)

We see that the y-terms have negative contribution to the energy of our system, and that energy has no lower bound (as well as no upper bound). For a usual physical

system this indicates its instability. But for a system derived from the Lagrangian (25) this is not the case. In eq.(25) the kinetic term for the y-component has negative sign, whilst that for the x-component has positive sign. Therefore, the equations of motion are

$$\ddot{x} = -\frac{\partial V}{\partial x} \,, \qquad \ddot{y} = \frac{\partial V}{\partial y} \tag{29}$$

where $V = \frac{1}{2}\omega^2(x^2 - y^2)$ is the potential. For the x-component the force is given by minus gradient of the potential, whilst for the y-component the force is given by plus gradient of the potential. Therefore, the criterium for the stability of motion for the y-degree of freedom is that the potential has to have a maximum in the (y, V)-plane. This is just opposite to the case of the x-degree of freedom where stability requires a minimum of V(x, y) in the plane (x, V).

This was just a more sophisticated explanation which involves the concept of energy. That our system is indeed stable can be directly read from the equations of motion (27) from which it follows that both x and y degrees of freedom oscillate around the origin x = 0, y = 0.

In the Hamiltonian form the equations of motion read

$$\dot{x} = \{x, H\} = \frac{\partial H}{\partial p_x} = p_x$$

$$\dot{y} = \{y, H\} = \frac{\partial H}{\partial p_y} = -p_y$$

$$\dot{p}_x = \{p_x, H\} = -\frac{\partial H}{\partial p_x} = -\omega^2 x$$

$$\dot{p}_y = \{p_y, H\} = -\frac{\partial H}{\partial p_y} = \omega^2 y$$
(30)

where the Poisson brackets are defined as usual. In particular we have

$${x, p_x} = 1, {y, p_y} = 1 (31)$$

The system can be quantized by replacing the canonically conjugate variables (x, p_x) and (y, p_y) by operators satisfying the following commutation relations⁷:

$$[x, p_x] = i$$
, $[y, p_y] = i$ (32)

Let us introduce non Hermitian operators

$$c_x = \frac{1}{\sqrt{2}} (\sqrt{\omega} x + \frac{i}{\sqrt{\omega}} p_x) , \quad c_x^{\dagger} = \frac{1}{\sqrt{2}} (\sqrt{\omega} x - \frac{i}{\sqrt{\omega}} p_x)$$
 (33)

⁷We use units in which $\hbar = c = 1$.

$$c_y = \frac{1}{\sqrt{2}} (\sqrt{\omega} y + \frac{i}{\sqrt{\omega}} p_y) , \quad c_y^{\dagger} = \frac{1}{\sqrt{2}} (\sqrt{\omega} y - \frac{i}{\sqrt{\omega}} p_y)$$
 (34)

They satisfy the commutation relations

$$[c_x, c_x^{\dagger}] = 1$$
, $[c_y, c_y^{\dagger}] = 1$ (35)

$$[c_x, c_y] = [c_x^{\dagger}, c_y^{\dagger}] = 0$$
 (36)

In terms of the new variables (33),(34) the Hamiltonian reads

$$H = \frac{1}{2}\omega(c_x^{\dagger}c_x + c_xc_x^{\dagger} - c_y^{\dagger}c_y - c_yc_y^{\dagger})$$
(37)

Let us define vacuum state according to

$$c_x|0\rangle = 0 , \qquad c_y|0\rangle = 0 \tag{38}$$

so that c_x , c_y are annihilation and c_x^{\dagger} , c_y^{\dagger} creation operators. Using (35) we find

$$H = \omega(c_x^{\dagger} c_x - c_y^{\dagger} c_y) \tag{39}$$

In the latter expression we have ordered the operators so that creation operators are on the left and annihilation operators on the right. We see that zero point energy in the Hamiltonian (39) cancels out!

It is instructive to consider the (x, y)-representation in which the momentum operators are $p_x = -i\partial/\partial x$, $p_y = -i\partial/\partial y$, and writing $\langle x, y | 0 \rangle \equiv \psi_0(x, y)$. Eqs. (38) then become

$$\frac{1}{2} \left(\sqrt{\omega} x + \frac{1}{\sqrt{\omega}} \frac{\partial}{\partial x} \right) \psi_0(x, y) = 0 \tag{40}$$

$$\frac{1}{2} \left(\sqrt{\omega} y + \frac{1}{\sqrt{\omega}} \frac{\partial}{\partial y} \right) \psi_0(x, y) = 0 \tag{41}$$

A solution which is in agreement with the probability interpretation reads

$$\psi_0 = \frac{2\pi}{\omega} e^{-\frac{1}{2}\omega(x^2 + y^2)} \tag{42}$$

and is normalized according to $\int \psi_0^2 dx dy = 1$. A particle described by the wave function ψ_0 of eq. (42) is localized around the origin. The excited states obtained by applying the product of operators c_x^{\dagger} , c_y^{\dagger} to the vacuum state are also localized. This is in agreement with the property that the corresponding classical particle moving

according to the equations of motion (27) is also "localized" in the vicinity of the origin.

All states $|\psi\rangle$ of our system have positive norm. For instance

$$\langle 0|cc^{\dagger}|0\rangle = \langle 0|[c,c^{\dagger}]|0\rangle = \langle 0|0\rangle = \int \psi^2 \,\mathrm{d}x \,\mathrm{d}y = 1 \tag{43}$$

This is so because of our choice of vacuum (38). Had we defined vacuum differently, e.g., by $c_x|0\rangle = 0$, $c_y^{\dagger}|0\rangle = 0$ we would have states with negative norm in our theory.

Our action (25) is invariant under the pseudo rotations in $M_{1,-1}$:

$$x' = \frac{x - vy}{\sqrt{1 - v^2}}, \qquad y' = \frac{y - vx}{\sqrt{1 - v^2}}$$
 (44)

where v is the parameter of the transformation. In a new reference frame S' we have $c_{x'}|0\rangle = 0$, $c_{y'}|0\rangle = 0$, and a normalized solution is

$$\psi_0' = \frac{2\pi}{\omega} e^{-\frac{\omega}{2}(x'^2 + y'^2)} \tag{45}$$

Expressed in terms of the old coordinates the latter wave function reads

$$\psi_0' = \frac{2\pi}{\omega} \exp\left[-\frac{\omega}{2}(x^2 + y^2)\frac{1 + v^2}{1 - v^2} + \frac{2\omega v}{1 - v^2}xy\right]$$
(46)

Let us now decide to observe everything from a fixed frame S. It is not difficult to find out that ψ'_0 of eq. (46) satisfies the Schrödinger equation in the old frame S:

$$-\frac{1}{2}\left(\frac{\partial^2 \psi_0'}{\partial x^2} - \frac{\partial^2 \psi_0'}{\partial y^2}\right) + \frac{\omega^2}{2}(x^2 - y^2)\psi_0' = 0 \tag{47}$$

In a given reference frame we have thus a family of solutions

$$\psi_0(x, y; v) = \frac{2\pi}{\omega} \exp\left[-\frac{\omega}{2}(x^2 + y^2)\frac{1 + v^2}{1 - v^2} + \frac{2\omega v}{1 - v^2}xy\right]$$
(48)

all having zero energy. For v = 0 we recover eq. (42).

The first excited state in the x and y direction, respectively, are

$$\psi_{10} = \frac{1}{\sqrt{2}} \left(\sqrt{\omega} x - \frac{1}{\sqrt{\omega}} \frac{\partial}{\partial x} \right) \psi_0 = \sqrt{2} \sqrt{\omega} x e^{-\frac{1}{2}\omega(x^2 + y^2)}$$
(49)

$$\psi_{01} = \frac{1}{\sqrt{2}} \left(\sqrt{\omega} y - \frac{1}{\sqrt{\omega}} \frac{\partial}{\partial y} \right) \psi_0 = \sqrt{2} \sqrt{\omega} y e^{-\frac{1}{2}\omega(x^2 + y^2)}$$
 (50)

In a new reference frame S' obtained by the transformation (44) the vacuum state is given by (45) and the excited states are given by the same equations (49),(50) in

which x and y are replaced by x' and y'. Expressing x' and y' in terms of x and y by using the transformation (44) we obtain

$$\psi_{10}(x,y;v) = \frac{\sqrt{2\omega}}{\sqrt{1-v^2}} (x-vy) \exp\left[-\frac{\omega}{2} \left(\frac{(x-vy)^2}{1-v^2} + \frac{(y-vx)^2}{1-v^2}\right)\right]$$
(51)

$$\psi_{01}(x,y;v) = \frac{\sqrt{2\omega}}{\sqrt{1-v^2}} (y-vx) \exp\left[-\frac{\omega}{2} \left(\frac{(x-vy)^2}{1-v^2} + \frac{(y-vx)^2}{1-v^2}\right)\right]$$
(52)

which are now the states as observed from the reference frame S. A state $\psi(x, y; v)$ is obtained from the state $\psi(x, y; 0)$ by an *active* pseudo rotation of the form (44).

One can verify that the states $\psi_{10}(x, y; v)$ and $\psi_{01}(x, y; v)$ satisfy the Schrödinger equation in frame S for arbitrary value of the parameter v:

$$\left[-\frac{1}{2} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} \right) + \frac{\omega}{2} \left(x^2 - y^2 \right) \right] \psi_{10}(x, y; v) = \omega \, \psi_{10}(x, y; v) \tag{53}$$

$$\left[-\frac{1}{2} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} \right) + \frac{\omega}{2} (x^2 - y^2) \right] \psi_{01}(x, y; v) = -\omega \,\psi_{01}(x, y; v) \tag{54}$$

The state $\psi_{10}(x, y; v)$ has energy $E = \omega$, whilst the state $\psi_{01}(x, y; v)$ has energy $E = -\omega$. A generic excited state $\psi_{nm}(x, y; v)$ has energy $E = \omega(n - m)$. Since v is an arbitrary parameter $v \in [0, 1]$ we have for fixed m, n a family of states $\{\psi_{mn}(x, y; v)\}$ all having the same energy $E = \omega(n - m)$. Not only the states (49),(50), but also the states (51),(52) and the higher excited states for arbitrary values of v satisfy the same Schrödinger equation.

A particular model of relativistic Harmonic oscillator in spacetime $M_{1,-3}$ has been considered in refs. [28] with a motivation to explain partons.

3.2 Generalization to $M_{r,-s}$

Let us now consider the harmonic oscillator in a pseudo-Euclidean space of arbitrary dimensiona and signature. Instead of (25) we have now the Lagrangian

$$L = \frac{1}{2}\dot{x}^{\mu}\dot{x}_{\mu} - \frac{1}{2}\omega^{2}x^{\mu}x_{\mu} \tag{55}$$

The canonical momenta are

$$p_{\mu} = \frac{\partial L}{\partial \dot{x}^{\mu}} = \dot{x}_{\mu} = \eta_{\mu\nu} \dot{x}^{\nu} \tag{56}$$

The Hamiltonia is

$$H = \frac{1}{2}p^{\mu}p_{\mu} + \frac{1}{2}\omega^{2}x^{\mu}x_{\mu} \tag{57}$$

Upon quantization the phase space variables (x^{μ}, p_{ν}) become operators satisfying

$$[x^{\mu}, p_{\nu}] = i\delta^{\mu}_{\ \nu} \quad \text{or} \quad [x^{\mu}, p^{\nu}] = i\eta^{\mu\nu}$$
 (58)

A straightforward generalization of the non Hermitian operators (33),(34) is

$$c^{\mu} = \frac{1}{\sqrt{2}} \left(\sqrt{\omega} x^{\mu} + \frac{i}{\sqrt{\omega}} p_{\mu} \right) , \qquad (59)$$

$$c^{\mu\dagger} = \frac{1}{\sqrt{2}} \left(\sqrt{\omega} x^{\mu} - \frac{i}{\sqrt{\omega}} p_{\mu} \right) \tag{60}$$

Notice that in the definition of c^{μ} , $c^{\mu\dagger}$ we take *contravariant* components of coordinates and *covariant* components of p_{μ} . This is in agreement with the definition (33),(34) where p_x , p_y defined in eq. (27) are in fact the covariant components. Using (59),(60), the Hamiltonian operator (57) becomes

$$H = \frac{1}{2}\omega(c^{\dagger}_{\mu}c^{\mu} + c^{\mu}c^{\dagger}_{\mu}) \tag{61}$$

The definition of vacuum state which is a generalization of the definition (38) reads

$$c^{\mu}|0\rangle = 0 \tag{62}$$

The annihilation and creation operators c^{μ} , $c^{\nu\dagger}$ satisfy the commutation relations

$$[c^{\mu}, c^{\nu\dagger}] = \delta^{\mu\nu} \tag{63}$$

where $\delta^{\mu\nu}$ is the Kronecker symbol (having +1 values on the diagonal and zero elsewhere). Using (63) we have

$$c^{\mu}c^{\dagger}_{\mu} = \eta_{\mu\nu}c^{\mu}c^{\nu\dagger} = \eta_{\mu\nu}(c^{\nu\dagger}c^{\mu} + \delta^{\mu\nu}) \equiv c^{\mu\dagger}c_{\mu} + r - s \tag{64}$$

where we have written $\eta_{\mu\nu}c^{\nu\dagger}c^{\mu} = \eta_{\mu\nu}c^{\mu\dagger}c^{\nu} \equiv c^{\mu\dagger}c_{\mu}$ and $\eta_{\mu\nu}\delta^{\mu\nu} = r - s$. Here r is the number of positive and s negative signature components.

The Hamiltonian is thus

$$H = \omega(c_{\mu}^{\dagger}c^{\mu} + \frac{r}{2} - \frac{s}{2}) \tag{65}$$

In particular, if r = s, the zero point energies cancel out.

The definition (59),(60) of the creation and annihilation operators is natural, because it takes a superposition of coordinates and canonical momenta. But from the viewpoint of the tensor calculus the notation in eq, (59),(60 is not very fortunate, because normally we do not sum covariant and contravariant components. Therefore I will now rewrite the theory by using the usual formalism in which creation and annihilation operators are defined in terms of contravariant components x^{μ} and p^{μ} :

$$a^{\mu} = \frac{1}{2} \left(\sqrt{\omega} \, x^{\mu} + \frac{i}{\sqrt{\omega}} \, p^{\mu} \right) \tag{66}$$

$$a^{\mu\dagger} = \frac{1}{2} \left(\sqrt{\omega} \, x^{\mu} - \frac{i}{\sqrt{\omega}} \, p^{\mu} \right) \tag{67}$$

The non vanishing commutators are

$$[a^{\mu}, a^{\dagger}_{\nu}] = \delta^{\mu}_{\ \nu} \quad \text{or} \quad [a^{\mu}, a^{\nu\dagger}] = \eta^{\mu\nu}$$
 (68)

Hamiltonian is

$$H = \frac{1}{2}\omega \left(a^{\mu\dagger} a_{\mu} + a_{\mu} a^{\mu\dagger} \right) \tag{69}$$

Given the operators (66),(67), let us consider two possible definitions of vacuum.

 1^{st} possible definition of vacuum

This is the definition that is usually adopted and it reads

$$a^{\mu}|0\rangle = 0 \tag{70}$$

The Hamiltonian, after using (68) and (70), is

$$H = \omega \left(a^{\mu \dagger} a_{\mu} + \frac{d}{2} \right) \tag{71}$$

where $d = \eta^{\mu\nu}\eta_{\mu\nu} = r + s$ is the dimension of $M_{r,-s}$. There occurs the non vanishing zero point energy. The eigenstates of H are all positive. This is so even for those terms in H which belong to negative signature: negative sign of a term in $a^{\mu\dagger}a_{\mu}$ is compensated by negative sign in the commutation relations (68). Also the expectation values between the eigenstates $|A\rangle$ of H calculated according to

$$\langle H \rangle = \frac{\langle A|H|A\rangle}{\langle A|A\rangle} \tag{72}$$

are always positive, since the negative norm in the denominator and negative norm in the numerator together give 1.

 2^{nd} possible definition of vacuum

Let us split $a^{\mu} = (a^{\alpha}, a^{\bar{\alpha}})$, where indices α , $\bar{\alpha}$ refer to the components with positive and negative signature, respectively, and define the vacuum according to [14] (see also [15]

$$a^{\alpha}|0\rangle = 0$$
 $a^{\bar{\alpha}\dagger}|0\rangle = 0$ (73)

Using (68) we obtain the Hamiltonian in which the annihilation operators, defined according to eq. (73), are on the right:

$$H = \omega \left(a^{\alpha \dagger} a_{\alpha} + \frac{r}{2} + a_{\bar{\alpha}} a^{\bar{\alpha}\dagger} - \frac{s}{2} \right) \tag{74}$$

where $\delta_{\alpha}{}^{\alpha} = r$ and $\delta_{\bar{\alpha}}{}^{\bar{\alpha}} = s$. If the number of positive and negative signature components is the same, i.e., r = s, then the Hamiltonian (74) has vanishing zero-point energy:

$$H = \omega (a^{\alpha \dagger} a_{\alpha} + a_{\bar{\alpha}} a^{\bar{\alpha} \dagger}) \tag{75}$$

Its eigenvalues are positive or negative, depending on which component (positive or negative signature) are excited. In x-representation the vacuum state (73) is

$$\psi_0 = \left(\frac{2\pi}{\omega}\right)^{d/2} \exp\left[-\frac{\omega}{2}\delta_{\mu\nu}x^{\mu}x^{\nu}\right] \tag{76}$$

where the Kronecker symbol $\delta_{\mu\nu}$ has values +1 for $\mu = \nu$ and 0 otherwise. It is a solution of the Schrödinger equation

$$-(1/2)\partial^{\mu}\partial_{\mu}\psi_{0} + (\omega^{2}/2)x^{\mu}x_{\mu}\psi_{0} = E_{0}\psi_{0}$$
(77)

with $E_0 = \omega(\frac{1}{2} + \frac{1}{2} + \dots - \frac{1}{2} - \frac{1}{2} - \dots)$. One can also easily verify that there are no negative norm states.

Let us now consider a pseudo rotation

$$x^{\prime \mu} = L^{\mu}_{\ \nu} x^{\nu} \tag{78}$$

The transformed vacuum wave function reads (see Sec.(3.1) for specific examples in $M_{1,-1}$):

$$\psi_0(x') = \exp\left[-\frac{\omega}{2}\,\delta_{\mu\nu}x'^{\mu}x'^{\nu}\right] = \exp\left[-\frac{\omega}{2}\delta_{\mu\nu}L^{\mu}{}_{\rho}L^{\nu}{}_{\sigma}x^{\rho}x^{\sigma}\right] = \psi'(x) \tag{79}$$

In order to show that also $\psi'(x)$ is a solution of the Schrödinger equation we use

$$\partial_{\beta}\psi_0' = -\omega \delta_{\mu\nu} L^{\mu}{}_{\beta} L^{\nu}{}_{\sigma} x^{\sigma} \psi' \tag{80}$$

and

$$\eta^{\alpha\beta}\partial_{\alpha}\partial_{\beta}\psi_{0}'(x) = \left(-\omega\eta^{\alpha\beta}L^{\mu}{}_{\alpha}L^{\nu}{}_{\beta}\delta_{\mu\nu} + \omega^{2}\eta^{\alpha\beta}\delta_{\mu\nu}L^{\mu}{}_{\beta}L^{\nu}{}_{\sigma}x^{\sigma}L^{\epsilon}{}_{\alpha}L^{\gamma}{}_{\rho}x^{\rho}\delta_{\epsilon\gamma}\right)\psi_{0}'(x)$$

$$= \omega^{2}x^{\alpha}x_{\alpha}\psi_{0}'(x) \tag{81}$$

where we have used $\eta^{\alpha\beta}L^{\mu}{}_{\beta}L^{\nu}{}_{\alpha} = \eta^{\mu\nu}$ and $\eta^{\mu\nu}\delta_{\mu\nu} = r - s = 0$. From eq.(81) it follows that $\psi'_0(x)$ satisfies

$$-\frac{1}{2}\eta^{\alpha\beta}\,\partial_{\alpha}\partial_{\beta}\psi_0'(x) + \frac{\omega^2}{2}\,x^{\alpha}x_{\alpha}\psi_0'(x) = 0 \tag{82}$$

which is the Schrödinger equation for the (actively) transformed vacuum wave function. Hence the theory is *covariant*, although the vacuum, defined according to eq. (73), is not *invariant* under the pseudo rotations.

In general, for the excited states, let us start from the Schrödinger equation in a reference frame S':

$$-\frac{1}{2}\eta^{\rho\sigma}\,\partial_{\rho}'\partial_{\sigma}'\psi(x')\frac{\omega^2}{2}\,x'^{\mu}x'_{\mu}\psi(x') = E\psi(x') \tag{83}$$

Let us now apply to eq.(83) the pseudo rotation $x'^{\mu} = L^{\mu}{}_{\rho}x^{\rho}$, briefly, x' = L(x):

$$-\frac{1}{2}\eta^{\rho\sigma} L^{\mu}{}_{\rho}L^{\nu}{}_{\sigma}\partial_{\mu}\partial_{\nu}\psi(L(x)) + \frac{\omega^2}{2} x^{\mu}x_{\mu}\psi(L(x))) = E\psi(L(x))$$
 (84)

Writing $\psi(L(x)) = \psi'(x)$ we find

$$-\frac{1}{2}\partial^{\mu}\partial_{\mu}\psi'(x) + \frac{\omega^2}{2}x^{\mu}x_{\mu}\psi'(x) = E\psi'(x)$$
(85)

which means that the (actively) transformed excited state $\psi'(x)$ (as observed from the frame S) satisfied the Schrödinger equation. So we have found that $\psi(x)$ as well as $\psi'(x)$ are solutions of the Schrödinger equation in S and they both have the same energy E. In general, in a given reference frame we have thus a degeneracy of solutions with the same energy (see also ref. [28]).

Quantum field theory 4

4.1 A system of scalar fields

Let us now turn to the theory of n scalar fields ϕ^a , a = 0, 1, 2, ..., n - 1, over the 4-dimensional spacetime parametrized by coordinates x^{μ} , $\mu = 0, 1, 2, 3$. The action for such system is

$$I[\phi^a] = \frac{1}{2} \int d^4x \sqrt{-g} (g^{\mu\nu} \partial_\mu \phi^a \partial_\nu \phi^b - m^2 \phi^a \phi^b) \gamma_{ab}$$
 (86)

where γ_{ab} is the metric in the space of fields ϕ^a , and $g_{\mu\nu}$ the metric of spacetime. Let us assume that $g_{\mu\nu} = \eta_{\mu\nu}$ is the metric of flat spacetime.

The canonical momenta are

$$\pi_a = \frac{\partial \mathcal{L}}{\partial \partial_0 \phi^a} = \partial^0 \phi_a = \partial_0 \phi_a \equiv \dot{\phi}_a \tag{87}$$

Upon quantization the following equal time commutation relations are satisfied:

$$[\phi^{a}(\mathbf{x}), \pi_{b}(\mathbf{x}')] = i\delta^{3}(\mathbf{x} - \mathbf{x}')\delta^{a}_{b}$$
(88)

The Hamiltonian is given by

$$H = \frac{1}{2} \int d^3x \left(\dot{\phi}^a \dot{\phi}^b - \partial_i \phi^a \partial^i \phi^b + m^2 \phi^a \phi^b \right) \gamma_{ab}$$
 (89)

where i = 1, 2, ..., n - 1. We shall assume that γ_{ab} is diagonal. A general solution to the equations of motion derived from the action (86) can be written in the form

$$\phi^{a} = \int \frac{\mathrm{d}^{3} \mathbf{k}}{(2\pi)^{3}} \frac{1}{2\omega_{\mathbf{k}}} (a^{a}(\mathbf{k}) \mathrm{e}^{-ikx} + a^{a\dagger}(\mathbf{k}) \mathrm{e}^{ikx})$$
(90)

Here⁸ $\omega_{\mathbf{k}} \equiv |\sqrt{m^2 + \mathbf{k}^2}|$. The creation and annihilation operators satisfy the commutation relations

$$[a^{a}(\mathbf{k}), a_{b}^{\dagger}(\mathbf{k}')] = (2\pi)^{3} 2\omega_{\mathbf{k}} \delta^{3}(\mathbf{k} - \mathbf{k}') \delta^{a}_{b}$$
(91)

or

$$[a^{a}(\mathbf{k}), a^{b^{\dagger}}(\mathbf{k}')] = (2\pi)^{3} 2\omega_{\mathbf{k}} \delta^{3}(\mathbf{k} - \mathbf{k}')\gamma^{ab}$$
(92)
⁸We use units in which $\hbar = c = 1$.

Inserting the expansion (90) of fields ϕ^a into the Hamiltonian (89) we obtain

$$H = \frac{1}{2} \int \frac{\mathrm{d}^3 \mathbf{k}}{(2\pi)^3} \frac{\omega_{\mathbf{k}}}{2\omega_{\mathbf{k}}} (a^{a\dagger}(\mathbf{k})a^b(\mathbf{k}) + a^a(\mathbf{k})a^{b\dagger}(\mathbf{k})) \gamma_{ab}$$
(93)

Let us assume that the signature of the metric γ_{ab} is pseudo-Euclidean, and let us write

$$a^{a}(\mathbf{k}) = (a^{\alpha}, a^{\bar{\alpha}}) \tag{94}$$

where α denotes positive and $\bar{\alpha}$ negative signature components.

We will define vacuum according to

$$a^{\alpha}(\mathbf{k})|0\rangle = 0 , \quad a^{\bar{\alpha}^{\dagger}}(\mathbf{k})|0\rangle = 0$$
 (95)

If we order the operators so that the annihilation operators, defined in eq. (95), are on the right, and use the commutation relations (92), we find

$$H = \int \frac{\mathrm{d}^3 \mathbf{k}}{(2\pi)^3} \frac{\omega_{\mathbf{k}}}{2\omega_{\mathbf{k}}} (a^{\alpha\dagger}(\mathbf{k}) a_{\alpha}(\mathbf{k}) + a^{\bar{\alpha}}(\mathbf{k}) a_{\bar{\alpha}}^{\dagger}) + \frac{1}{2} \int \mathrm{d}^3 \mathbf{k} \, \omega_{\mathbf{k}} \delta^3(0) (r - s) \tag{96}$$

where $r = \delta^{\alpha}{}_{\alpha}$ and $s = \delta^{\bar{\alpha}}{}_{\bar{\alpha}}$. In the case in which the signature has equal number of plus and minus signs, i.e., when r = s, the zero point energies cancel out from the Hamiltonian.

4.2 Genaralization to Clifford space

Now a question arises as to why should the space of fields have the metric with r=s. Isn't it an ad hoc assumption? The answer is as follows. We can consider the space of fields V_n just as a starting space, with basis e_a , a=0,1,2,...,n-1, from which we generate the 2^n -dimensional Clifford space C_{V_n} with basis $e_A=(\underline{1},e_a,e_{a_1a_2},...,e_{a_1...a_n})$, $a_1 < a_2 < ... < a_n$. If V_n is a Euclidean space so that $e_a \cdot e_b = \delta_{ab}$ is the Euclidean metric, then also the metric $e_A^{\dagger} * e_B$ of C_{V_n} is Euclidean. But, as it was pointed out in refs. [9, 6], instead of the basis e_A we can take another basis, e.g.,

$$\gamma_A = (\underline{1}, \gamma_a, \gamma_{a_1 a_2}, \dots, \gamma_{a_1 \dots a_n}) \tag{97}$$

generated from the set of Clifford numbers $\gamma_a=(e_0,e_ie_0),\ a=0,1,2,...,n-1;\ i=1,2,...,n$ satisfying

$$\gamma_a \cdot \gamma_b \equiv \frac{1}{2} (\gamma_a \gamma_b + \gamma_b \gamma_a) = \eta_{ab} \tag{98}$$

The metric

$$\gamma_A^{\dagger} * \gamma_B = G_{AB} \tag{99}$$

defined with respect to the new basis is pseudo-Euclidean, its signature having $2^n/2$ plus and $2^n/2$ minus signs.

We assume that a field theory should be formulated in C-space in which the metric is given by eq.(99). Instead of the action (86) we thus consider its generalization to C-space:

$$I = \frac{1}{2} \int d^4x \sqrt{-g} (g^{\mu\nu} \partial_{\mu} \phi^A \partial_{\nu} \phi^B - m^2 \phi^A \phi^B) G_{AB}$$
 (100)

Here $\phi^A \gamma_A$ is a polyvector field. Since the metric G_{AB} has signature (+ + + ... - - ...) = (R+, S-) with R = S, zero point energies of a system based on the action (100) cancel out: vacuum energy vanishes. Consequently, in such a theory there is no cosmological problem [15]. The small cosmological constant, as recently observed, could be a residual effect of something else.

Cancellation of vacuum energies in the theory does not exclude [15] the existence of well known effects, such as Casimir effect, which is a manifestation of vacuum energies.

3. Strings and Clifford space

Usual strings are described by the mapping $(\tau, \sigma) \to x^{\mu} = X^{\mu}(\tau, \sigma)$, where the embedding functions $X^{\mu}(\tau, \sigma)$ describe a 2-dimensional worldsheet swept by a string. The action is given by the requirement that the area of the worldsheet be "minimal" (extremal). Such action is invariant under reparametrizations of (τ, σ) . There are several equivalent forms of the action including the " σ -model action" which, in the conformal gauge, can be written as

$$I[X^{\mu}] = \frac{\kappa}{2} \int d\tau \, d\sigma \, (\dot{X}^{\mu} \dot{X}_{\mu} - X^{\prime \mu} X_{\mu}^{\prime}) \tag{101}$$

where $\dot{X}^{\mu} \equiv \mathrm{d}X^{\mu}/\mathrm{d}\tau$ and $X'^{\mu} \equiv \mathrm{d}X^{\mu}/\mathrm{d}\sigma$. Here κ is the string tension, usually written as $\kappa = 1/(2\pi\alpha')$.

String coordinates X^{μ} and momenta $P_{\mu} = \partial L/\partial \dot{X}^{\mu} = \kappa \dot{X}_{\mu}$ satisfy the following constraints $(\sigma \in [0, \pi])$:

$$\varphi_1(\sigma) = P^{\mu} P_{\mu} + \frac{X'^{\mu} X'_{\mu}}{(2\pi\alpha')^2} \approx 0 \qquad \varphi_2(\sigma) = \frac{P^{\mu} X'_{\mu}}{\pi\alpha'} \approx 0 \qquad (102)$$

which can be written as a single constraint on the interval $\sigma \in [-\pi, \pi]$

$$\Pi^{\mu}\Pi_{\mu}(\sigma) \approx 0 \qquad \Pi^{\mu} = P^{\mu} + \frac{X^{\prime \mu}}{2\pi\alpha^{\prime}}$$
 (103)

to which the open string is symmetrically extended. (For more details see the literature on strings, e.g., [29].)

If we generalize the action (101) to C-space, we obtain

$$I[X] = \frac{\kappa}{2} \int d\tau \, d\sigma \, (\dot{X}^M \dot{X}^N - X'^M X'_N) G_{MN}$$
 (104)

where κ is the generalized string tension. Taking 4-dimensional spacetime, there are $D=2^4=16$ dimensions of the corresponding C-space. Its signature (+ + +... - - - ...) has 8 plus and 8 minus signs. The variables X^M are components of a polyvector X expanded according to eq. (22) and they depend on two parameters τ and σ . From the point of view of C-space the variables $X^M(\tau,\sigma)$ describe an object with two intrinsic dimensions, that is, a 2-dimensional 'world sheet' living in a 16-dimensional C-space. Therefore we will keep on talking about 'strings' (that sweep a world sheet).

Let us consider the case of an open string satisfying the boundary condition $X'^{M} = 0$ at $\sigma = 0$ and $\sigma = \pi$. Then we can make the expansion

$$X^{M}(\tau,\sigma) = \sum_{n=-\infty}^{\infty} X_{n}^{M}(\tau) e^{in\sigma}$$
(105)

where from the reality condition $(X^M)^* = X^M$ it follows

$$X_n^M = X_{-n}^M (106)$$

Inserting (105) into (104), integrating over σ and taking into account (106) we obtain the action expressed in terms of $X_n^M(\tau)$:

$$I[X_n^M] = \frac{\kappa'}{2} \int d\tau \sum_{n=-\infty}^{\infty} (\dot{X}_n^M \dot{X}_n^N - n^2 X_n^M X_n^N) G_{MN}$$
 (107)

where $\kappa' = 2\pi\kappa = 1/\alpha'$. This is just the action of infinite number of harmonic oscillators.

The Hamiltonian corresponding to the action (107) is

$$H = \frac{1}{2} \sum_{n=-\infty}^{\infty} \left(\frac{1}{\kappa'} P_n^M P_{nM} + \kappa' \, n^2 X_n^M X_{nM} \right) \tag{108}$$

Let us introduce

$$a_n^M = \frac{1}{\sqrt{2}} \left(\frac{1}{\sqrt{\kappa'}} P_n^M - in\sqrt{\kappa'} X_n^M \right)$$

$$a_n^{M\dagger} = \frac{1}{\sqrt{2}} \left(\frac{1}{\sqrt{\kappa'}} P_n^M + in\sqrt{\kappa'} X_n \right)$$
(109)

We see that

$$a_{-n}^M = a_n^{M\dagger} \tag{110}$$

Rewriting H in terms of a_n^M , $a_n^{M\dagger}$ we obtain

$$H = \frac{1}{2} \sum_{n=-\infty}^{\infty} (a_n^{M\dagger} a_{nM} + a_{nM} a_n^{M\dagger}) = \sum_{n=1}^{\infty} (a_n^{M\dagger} a_{nM} + a_{nM} a_n^{M\dagger}) + \frac{1}{2\kappa'} P_0^M P_{0M}$$
 (111)

Upon quantization we have

$$[X_n^M, P_{nN}] = i\delta^M{}_N \quad \text{or} \quad [X_n^M, P_n^N] = iG^{MN}$$
 (112)

and

$$[a_n^M, a_{nN}^{\dagger}] = n\delta^M{}_N \quad \text{or} \quad [a_n^M, a_n^{N\dagger}] = nG^{MN}$$
 (113)

In order to construct the Fock space of excited states, one has first to define a vacuum state. There are two possible choices [15].

Possibility I. Conventionally, the vacuum state is defined according to

$$a_n^M|0\rangle = 0 , \quad n \ge 1 \tag{114}$$

and the excited part of the Hamiltonian $H_{\text{exc}} = H - (1/\kappa')P_0^M P_{0M}$, after using (113) and (114) is

$$H_{\text{exc}} = \sum_{n=-\infty}^{\infty} (a_n^{M\dagger} a_{nM} + \frac{n}{2}D) = 2\sum_{n=1}^{\infty} (a_n^{M\dagger} a_{nM} + \frac{n}{2}D)$$
 (115)

$$D = \delta^M{}_M = G^{MN}G_{MN}$$

Its eigenvalues are all positive⁹ and there is the non vanishing zero point energy. But there exist negative norm states.

⁹This is so even for those components a_n^M that belong to negative signature: negative sign of a term in $a_n^{M\dagger}a_{nM}$ is compensated by negative sign in the commutation relation (113).

Possibility II. Let us split $a_n^M = (a_n^A, a_n^{\bar{A}})$ where the indices A, \bar{A} refer to the components with positive and negative signature, respectively, and let us define vacuum according to

$$a_n^A|0\rangle = 0 , \qquad (a_n^{\bar{A}})^{\dagger}|0\rangle = 0 , \quad n \ge 1$$
 (116)

Using (113) we obtain the Hamiltonian in which the annihilation operators, defined according to eq.(116), are on the right:

$$H_{\text{exc}} = 2\sum_{n=1}^{\infty} \left(a_n^{A\dagger} a_{nA} + \frac{n}{2} D_+ + a_{n\bar{A}} a_n^{\bar{A}\dagger} - \frac{n}{2} D_-\right)$$
 (117)

where

$$D_{+} = \delta_A^A , \qquad D_{-} = \delta_{\bar{A}}^{\bar{A}} \tag{118}$$

are, respectively, the number of positive and negative signature dimensions and $D = D_+ + D_- = \delta_M^M$ the total number of dimensions of Clifford space. There are no negative norm states.

Since in Clifford space the number of positive and negative signature components is the same, i.e., $D_{+} = D_{-}$, the above Hamiltonian has vanishing zero point energy:

$$H_{\text{exc}} = 2\sum_{n=1}^{\infty} (a_n^{A^{\dagger}} a_{nA} + a_{n\bar{A}} a_n^{\bar{A}^{\dagger}})$$
 (119)

Its eigenvalues can be positive or negative, depending on which components (positive or negative signature) are excited.

An immediate objection could arise at this point, namely, that since the spectrum of the Hamiltonian is not bounded from below, the system described by H of eq.(117) or (119) is unstable. This objection would only hold if the kinetic terms $\dot{X}_n^M \dot{X}_{nM}$ in the action (107) (or the terms $P_n^M P_{nM}$ in the Hamiltonian (108)) were all positive, so that negative eigenvalues of H would come from the negative potential terms in $n^2 X_n^M X_{nM}$. But since our metric is pseudo-Euclidean, whenever a term in the potential is negative, also the corresponding kinetic term is negative. Therefore, the acceleration corresponding to negative signature term is proportional to the plus gradient of potential (and not to the minus gradient of potential as it is the case for positive signature term); such system is stable if the potential has maximum, i.e., if it has an upper bound (and not a lower bound). The overall change of sign of

the action (Lagrangian) has no influence on the equations of motion (and thus on stability).

In the bosonic string theory based on the ordinary definition of vacuum (Possibility I) and formulated in D-dimensional spacetime with signature (+-----) there are negative norm states, unless D=26. Consistency of the string theory requires extra dimensions, besides the usual four dimensions of spacetime.

My proposal is that, instead of adding extra dimensions to spacetime, we can start from 4-dimensional spacetime M_4 with signature (+---) and consider the Clifford space C_{M_4} (C-space) whose dimension is 16 and signature (8+,8-). The necessary extra dimensions for consistency of string theory are in C-space. This also automatically brings spinors into the game. It is an old observation that spinors are the elements of left or right ideals of Clifford algebras [16]–[18] (see also a very lucid and systematic recent exposition in refs. [25]). In other words, spinors are particular sort of polyvectors [6]. Therefore, the string coordinate polyvectors contain spinors. This is an alternative way of introducing spinors into the string theory [6, 11]. An attempt to achieve a deeper understanding of the structure of supersymmetry has beeb undertaken in refs. [30].

Let the constraints (102), (103) be generalized to C-space. So we obtain

$$\Pi^M \Pi_M \approx 0 , \qquad \Pi^M = P^M + \frac{X'^M}{2\pi\alpha'}$$
 (120)

Using (105) and expanding the momentum $P^{M}(\sigma)$ according to

$$P^{M} = \sum_{n=-\infty}^{\infty} P_{n}^{M} e^{in\sigma}$$
 (121)

we can calculate the Fourier coefficients of the constraint (called Virasoro generators):

$$L_n = \frac{\pi \alpha'}{2} \int_{-\pi}^{\pi} d\sigma \, e^{-in\sigma} \, \Pi^M \Pi_M = \frac{1}{2} \sum_{r=-\infty}^{\infty} a_{-r}^M a_{r-n}^N G_{MN}$$
 (122)

Let us now calculate the commutators of Virasoro generators. If $m \neq -n$ the commutators can be straightforwardly calculated. The result is

$$[L_n, L_m] = \frac{1}{2}(n-m)\sum_{r=-\infty}^{\infty} a_r^M a_{(n+m-r)M} = (n-m)L_{n+m} , \quad m \neq -n$$
 (123)

For m = -n we have to bear in mind that a_{-r}^M and a_r^M due to eqs.(110) and (113) do not commute. Therefore, if we put the operators which annihilate the vacuum

according to eq.(116) on the right, we obtain the following expression:

$$[L_{n}, L_{-n}] = n \sum_{r=-\infty}^{\infty} a_{r}^{M} a_{-rM}$$

$$= n \left(a_{0}^{M} a_{0M} + \sum_{r=1}^{\infty} (a_{-r}^{A} a_{rA} + a_{-r}^{\bar{A}} a_{r\bar{A}} + a_{r}^{A} a_{-rA} + a_{r}^{\bar{A}} a_{-r\bar{A}}) \right)$$

$$= n \left(a_{0}^{M} a_{0M} + \sum_{r=1}^{\infty} (2a_{-r}^{A} a_{rA} + rD_{+} + 2a_{r}^{\bar{A}} a_{-r\bar{A}} - rD_{-}) \right)$$

$$= 2nL_{0} + n \sum_{r=1}^{\infty} r(D_{+} - D_{-}) = 2nL_{0}, \quad m = -n$$
(124)

where in the last step we have taken into account that in C-space the number of positive and negative signature dimensions is the same, i.e., $D_+ = D_-$.

Combining together eqs. (123) and (124) we find the following relation

$$[L_m, L_n] = (m - n)L_{n+m} (125)$$

which holds for arbitrary positive or negative integers n and m. There are no central terms. The terms which arise after normal ordering the operators in eq. (124) have opposite sign for positive and negative signature components, and thus cancel out. The algebra of Virasoro generators is thus closed, which automatically assures consistency of quantum string theory formulated in 16-dimensional Clifford space generated by the spacetime vectors γ_{μ} .

5 Conclusion

We have shown that generalizing physics from 4-dimensional spacetime to Clifford space (C-space) has promising consequences for quantum field theory and string theory. In order to avoid inconsistencies, the definition of C-space metric as the scalar product of two basis elements has to involve reversion. The metric so defined has signature ($+ + + \dots - - \dots$) (eight times plus, eight times minus). In QFT formulated in C-space the vacuum energy vanishes, if the vacuum state is defined in a very natural and straightforward way as proposed by Jackiw [14] (see also [15]). Therefore, there is no cosmological constant problem in such a theory. Generalizing the 4-dimensional target space (in which a string lives) to a 16-dimensional Clifford space, we have found that the (quantum) algebra of Virasoro generators has no

central terms. String theory is consistent if formulated in 16-dimensional Clifford space. This automatically brings fermions into the game, since fermions are the elements of left and right ideals of a Clifford algebra. Therefore the C-space formulation of string theory is an alternative to the usual superstring formulation which involves 10 extra dimensions of target space and Grassmann odd variables. While in the usual string theory one has to go beyond the 4-dimensional spacetime, and then study how to compactify the unobservable extra dimensions, in the proposed new theory we remain in 4-dimensional spacetime. The "extra dimensions" reside in Clifford space, and they have a physical interpretation as providing a description of extended object [10, 6]. We have thus found a very promising outline of QFT and string theory which deserves further studies.

References

- [1] P.A.M. Dirac, talk presented at International School of Subnuclear Physics; 19th Course: The Unity of the Fundamental Interactions, 31 July–11 August 1981, Erice, Sicily, Italy.
- [2] R. P. Feynman, Phys. Rev. 84, 108 (1951).
- [3] S.S. Schweber, Rev. Mod. Phys. 58, 449 (1986).
- [4] V. Fock, Phys. Z. Sowj. 12, 404 (1937). E.C.G. Stueckelberg, Helv. Phys. Acta,
 14, 322 (1941); 14, 588 (1941); 15, 23 (1942).
- [5] L. P. Horwitz and C. Piron, Helv. Phys. Acta 46, 316 (1973). L. P. Horwitz and F. Rohrlich, Physical Review D 24, 1528 (1981); 26, 3452 (1982). L. P. Horwitz, R. I. Arshansky and A. C. Elitzur Found. Phys 18, 1159 (1988). R. Arshansky, L. P. Horwitz and Y. Lavie, Foundations of Physics 13, 1167 (1983). L. P. Horwitz, in Old and New Questions in Physics, Cosmology, Philosophy and Theoretical Biology (Editor Alwyn van der Merwe, Plenum, New York, 1983). L. P. Horwitz and Y. Lavie, Physical Review D 26, 819 (1982). L. Burakovsky, L. P. Horwitz and W. C. Schieve, Physical Review D 54, 4029 (1996). L. P. Horwitz and W. C. Schieve, Annals of Physics 137, 306 (1981). J.R.Fanchi, Phys. Rev. D 20, 3108 (1979). See also the review J.R.Fanchi, Found. Phys. 23, 287 (1993),

- and many references therein; J. R. Fanchi *Parametrized Relativistic Quantum Theory* (Kluwer Academic, Dordrecht, 1993). M. Pavšič, *Found. Phys.* **21**, 1005 (1991); M. Pavšič, *Nuovo Cim.* **A104**, 1337 (1991).
- [6] M.Pavšič: The Landscape of Theoretical Physics: A Global View; From Point Particle to the Brane World and Beyond, in Search of Unifying Principle, Kluwer Academic, Dordrecht 2001.
- [7] W. Pezzaglia, Physical Applications of a Generalized Geometric Calculus [arXiv: gr-qc/9710027]; Dimensionally Democratic calculus and Principles of Polydimensional Physics [arXiv: gr-qc/9912025]; Classification of Multivector Theories and Modifications of the Postulates of Physics [arXiv: gr-qc/9306006]; Physical Applications of Generalized Clifford Calculus: Papatetrou equations and Metamorphic Curvature [arXiv: gr-qc/9710027]; Classification of Multivector theories and modification of the postulates of Physics [arXiv: gr-qc/9306006].
- [8] C. Castro, Chaos, Solitons and Fractals 10 (1999) 295; Chaos, Solitons and Fractals 12 (2001) 1585; The Search for the Origins of M Theory: Loop Quantum Mechanics, Loops/Strings and Bulk/Boundary Dualities [arXiv: hep-th/9809102]; C. Castro, Chaos, Solitons and Fractals 11 (2000) 1663; Foundations of Physics 30 (2000) 1301.
- [9] M. Pavšič, Found. Phys. **31** (2001) 1185 [arXiv:hep-th/0011216].
- $[10]\ \mathrm{M.\ Pav\check{s}i\check{c}},\ Found.\ Phys.\ \mathbf{33}\ (2003)\ 1277\ [arXiv:gr-qc/0211085].$
- [11] C. Castro and M. Pavšič, *Phys. Lett. B* $\bf 539$ (2002) 133 [arXiv:hep-th/0110079].
- [12] A. Aurilia, S. Ansoldi and E. Spallucci, Class. Quant. Grav. 19 (2002) 3207 [arXiv:hep-th/0205028].
- [13] C. Castro and M. Pavšič, *Int. J. Theor. Phys.* **42** (2003) 1693 [arXiv:hep-th/0203194].
- [14] D. Cangemi, R. Jackiw and B. Zwiebach, Annals of Physics 245 (1996) 408; E. Benedict, R. Jackiw and H.-J. Lee, Phys. Rev. D 54 (1996) 6213

- [15] M. Pavšič, *Phys. Lett. A* **254** (1999) 119 [arXiv:hep-th/9812123].
- [16] S. Teitler, Supplemento al Nuovo Cimento III, 1 (1965) and references therein;
 Supplemento al Nuovo Cimento III, 15 (1965); Journal of Mathematical Physics
 7, 1730 (1966); Journal of Mathematical Physics 7, 1739 (1966). L. P. Horwitz,
 J. Math. Phys. 20, 269 (1979); H. H. Goldstine and L. P. Horwitz, Mathematische Annalen 164, 291 (1966).
- [17] D. Hestenes, Space-Time Algebra, Gordon and Breach, New York, 1966; D. Hestenes and G. Sobcyk, Clifford Algebra to Geometric Calculus, D. Reidel, Dordrecht, 1984.
- [18] P. Lounesto, *Clifford Algebras and Spinors* (Cambridge University Press, Cambridge, 2001).
- [19] W. Baylis, *Electrodynamics*, A Modern Geometric Approach (Boston, Birkhauser, 1999).
- [20] G. Trayling and W. Baylis, J.Phys. A 34, 3309 (2001); Int.J.Mod.Phys. A 16S1C, 909 (2001).
- [21] B. Jancewicz, Multivectors and Clifford Algebra in Electrodynamics (World Scientific, Singapore 1989).
- [22] Clifford Algebras and their applications in Mathematical Physics Vol 1: Algebras and Physics. eds by R. Ablamowicz, B. Fauser. Vol 2: Clifford analysis. eds by J. Ryan, W. Sprosig Birkhauser, Boston 2000;
- [23] A. Lasenby and C. Doran, *Geometric Algebra for Physicists* (Cambridge U. Press, Cambridge 2002).
- [24] A.M.Moya, V.V Fernandez and W.A. Rodrigues, Int.J.Theor.Phys. 40 (2001) 2347-2378 [arXiv: math-ph/0302007]; Multivector Functions of a Multivector Variable [arXiv: math.GM/0212223]; Multivector Functionals [arXiv: math.GM/0212224]; W.A. Rodrigues, Jr, J. Vaz, Jr, [Adv. Appl. Clifford Algebras 7 (1997) 457-466. E.C de Oliveira and W.A. Rodrigues, Jr, Ann. der

- Physik 7 654-659 (1998); Phys. Lett. A 291 367 (2001); W.A. Rodrigues, Jr.
 J.Y.Lu, Foundations of Physics 27 (1997) 435-508.
- [25] N. S. Mankoč Borštnik and H. B. Nielsen, *J. Math. Phys.* **43** 5782 (2002) [arXiv:hep-th/0111257]; *J. Math. Phys.* **44** 4817 (2003) [arXiv:hep-th/0303224].
- [26] See e.g., S. Weinberg, Rev. Mod. Phys. **61**, 1 (1989).
- [27] S. Perlmutter, et al., Astrophys. J. 517, 565 (1999); A.G. Riess, et al., Astron.
 J. 116, 1009 (1998); D. N. Spergel, et al., Astrophys. J. Suppl 148, 175 (2003);
 L. Page, et al., Astrophys. J. Suppl. 148, 233 (2003);
- [28] Y.S. Kim and M.E. Noz, Phys. Rev. D 8, 3521 (1973); 12, 122 (1975); 15,335 (1977); Phys. Rev. Lett. 63, 348 (1989); Y.S. Kim and M.E. Noz, Theory and Applications of the Poincareé Group (D. Reidel Publishing Company, Dordrecht, 1986).
- [29] See e.g., M.B. Green, J.H. Schwarz and E. Witten, Superstring Theory, Cambridge University Press, 1987; M. Kaku: Introduction to Superstrings, Springer-Verlag, New York, N.Y., 1988.
- [30] M. Faux and S. J. Gates, Adinkras: A graphical technology for supersymmetric representation theory, arXiv:hep-th/0408004; S. J. J. Gates, W. D. Linch and J. Phillips, When superspace is not enough, arXiv:hep-th/0211034; S. J. J. Gates and L. Rana, Phys. Lett. B 369, 262 (1996) [arXiv:hep-th/9510151]; S. J. Gates and L. Rana, Phys. Lett. B 352, 50 (1995) [arXiv:hep-th/9504025].